# GRAVITATIONAL RADIATION AND SPIRALLING TIME OF CLOSE BINARY SYSTEMS (III)

(Letter to the Editor)

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**Abstract.** Twenty-seven typical binary systems have been investigated to study their power-output by gravitational radiation  $(P_B)$  and spiralling time  $(\tau_0)$  relation. It was found that these binary systems form two distinct groups. New relations have been given between  $P_B$  and  $\tau_0$  for each group.

## 1. Introduction

In continuation of our (Padalia, 1987, 1988) earlier work on this subject we have evaluated  $P_B$  and  $\tau_0$  values for 27 close binary systems for which absolute dimensions are known. New relations between  $P_B$  and  $\tau_0$  have been established. Out of these 27 systems, 19 binaries appear to form one group since they lie along one curve whereas the remaining eight systems follow another curve. The systems considered, for gravitational radiation study, are in the mass range  $0.5 M_{\odot}$  to  $47 M_{\odot}$ . Unlike our second paper (Padalia, 1988) where almost all the systems discussed were B-type, the spectral type of the systems described in the present paper are from O to K.

The masses, period, spectral type, and radii of relative orbits adopted in the present paper are given in Table I.

# 2. Discussions and Results

For determining  $P_B$  and  $\tau_0$  values of individual systems, equations and assumptions used by us have been reported in Paper I (Padalia, 1987). However, for the sake of completeness, the equations are

$$P_B = \left(\frac{\mu}{M_\odot}\right)^2 \left(\frac{M}{M_\odot}\right)^{4/3} P^{-10/3} = 3.0 \times 10^{26} \,\mathrm{W}\,,\tag{1}$$

$$\tau_0 = \frac{5c^5 a_0^4}{256G^3 \mu M^2} \ . \tag{2}$$

The values of  $P_B$  and  $\tau_0$  have been determined for 27 binary systems. Gravitational radiation  $P_B$  (in watts) along the X-axis and spiralling time  $\tau_0$  (in years) along the Y-axis have been plotted in Figure 1. It is found that X and Y follow the relations

$$Y = 0.0071X^3 + 0.1X^2 - 1.18X + 3.30$$

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 $\label{eq:TABLE} TABLE\ I(a)$  Gravitational radiation and spiralling time of nineteen binary systems

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Name of the binary systems	Sp. type	$M_1(M_\odot)$	$M_2(M_\odot)$	Period in days	Radius of relative orbit $a_0(R_{\odot})$	Power output $(P_B X)$ (W)	Spiral time $(\tau_0)$ (years)	$X \tag{log} P_B - 19)$	$Y$ (log $ au_0-10$ )
AO Cas	O9III, O9III	23.0	18.0	3.524	32.77	$162858.4 \times 10^{19}$	$1.01 \times 10^{10}$	5.218	0.004
V444 Cyg	06, WN54.5	25.6	10.1	4.212	35.85	$38428.4 \times 10^{19}$	$2.06 \times 10^{10}$	4.584	0.315
CW Cep	B0.5, B0.5IV	11.8	11.1	2.730	23.30	$34467.5 \times 10^{19}$	$1.898 \times 10^{10}$	4.537	0.276
AO Mon	B3, B5	5.5	5.2	1.885	14.17	$15712.5 \times 10^{19}$	$1.92 \times 10^{10}$	4.196	0.283
U Oph	B5V, B5V	5.0	4.5	1.677	12.85	$15369.6 \times 10^{19}$	$1.87 \times 10^{10}$	4.186	0.273
U Her	M6.5E, G8III-IV	8.0	2.8	2.050	13.60	$7154.0 \times 10^{19}$	$2.08 \times 10^{10}$	3.854	0.319
AV Vir	K0V, K0V	1.4	9.0	0.408	2.22	$6106.7 \times 10^{19}$	$2.30 \times 10^{10}$	3.785	0.362
RX Her	A0V, A0V	2.7	2.3	1.779	9.74	$1494.2 \times 10^{19}$	$4.21\times10^{10}$	3.174	0.624
TV Cas	B9V, F7IV	3.1	1.4	1.813	8.89	$706.5 \times 10^{19}$	$4.80 \times 10^{10}$	2.849	0.681
CM Lac	A2V, A8V	2.0	1.5	1.605	8.51	$610.9 \times 10^{19}$	$7.38 \times 10^{10}$	2.786	898.0
U Cep	B7VE, G8III-IV	3.1	1.5	2.493	12.62	$302.9 \times 10^{19}$	$14.88 \times 10^{10}$	2.481	1.214
YY Gem	dM, E, –	0.5	0.5	0.814	5.02	$136.2 \times 10^{19}$	$27.02 \times 10^{10}$	2.134	1.432
WW Dra	G2IV, K0IV	4.0	2.3	4.630	21.60	$121.2 \times 10^{19}$	$53.52 \times 10^{10}$	2.083	1.726
TX Leo	A2V, –	2.4	1.0	2.445	11.10	$97.0 \times 10^{19}$	$27.6 \times 10^{10}$	1.987	1.441
$\beta$ Aur	AP, B9V	2.3	2.2	3.960	19.57	$76.7 \times 10^{19}$	$90.3 \times 10^{10}$	1.884	1.955
AR Aur	AP, B9V	2.5	2.3	4.135	18.47	$75.4 \times 10^{19}$	$63.88 \times 10^{10}$	1.877	1.805
TX UMa	B8V, G0III	3.1	1.9	3.063	13.76	$56.5 \times 10^{19}$	$47.07 \times 10^{10}$	1.753	1.672
$\beta$ Per	G2IB	3.1	0.7	2.867	11.71	$49.3 \times 10^{19}$	$32.00 \times 10^{10}$	1.693	1.505
Z Her	FIV-V	1.2	1.1	3.993	15.09	$7.6 \times 10^{19}$	$14.88 \times 10^{10}$	0.883	2.394

 $\label{eq:TABLE} TABLE\ I(b)$  Gravitational radiation and spiralling time of eight binary systems

in relative days orbit $a_0(R_{\odot})$ 4.393 53.18 1.775 18.05	37.2
	1.874
2.21	0.332
1.15	0.278
27.66	12.426
2.836	0.339

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for 19 binary systems and

$$Y = 0.23X^2 - 1.889X + 2.84$$

for remaining 8 binary systems, where

$$X = \log P_B - 19$$
 and  $Y = \log \tau_0 - 10$ .

The curves representing these equations are shown by solid lines in Figure 1. It is interesting to note that eight systems lying along the curve II in Figure 1 actually have the spiralling time range  $10^9$  year whereas the upper curve I in the same figure gives a spiralling time range of  $10^{10}$  years for the remaining 19 binaries. The binaries lying along curves I and II of Figure 1 are listed in Tables I(a) and I(b), respectively. In our earlier papers we have found that  $P_B$  is inversely proportional to  $\tau_0$ . This finding is also true for almost 19 binary systems of Table I(a). However, this relation of  $P_B$  versus  $\tau_0$  is not true for the remaining 8 systems listed in Table I(b).

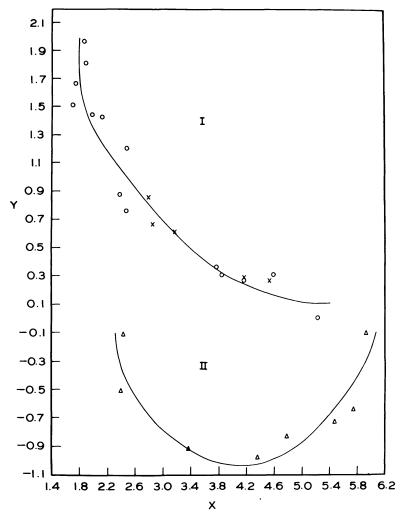


Fig. 1. Relation between spiralling time and gravitational radiation for 27 binary systems, curve I stands for 19 binary systems given in Table I(a) and curve II stands for the remaining 8 systems listed in Table I(b).

In Figure 1 of Paper I it was mentioned that the binary systems lie in two different parts of the curve, one from A to B and another from B' to C. The 5 systems lying in B'C part of the curve were CM Lac, TV Cas, RX Her, AO Mon, and DW Cep. This left a question whether B'C is the part of the curve AB or a separate curve. Here we find that all these 5 systems, making mostly B'C part of the earlier curve, fit very well in the present curve I of Figure 1. This indicates that these 5 systems belong to the present curve I of Figure 1 but not of Figure 1 of Paper I. These 5 systems are marked as crosses in Figure 1.

It is concluded from the present study that a large number of binary systems can be arranged to follow a few equations (of  $P_B$  and  $\tau_0$  relations), and subsequently an individual system can be catalogued as Equation (1) or Equation (2) binary, etc. For example in our Paper II (Padalia, 1988) all binaries in the mass range 7 to 19  $M_{\odot}$  and  $P_B$ ,  $\tau_0$  range to  $10^{23}$  W and  $10^9$  years, respectively, were found to be early type, mostly B-type binaries. However, this could not be true for the present 27 systems. For conclusive results investigation of a large number of binary systems is needed for finding out common properties of the systems following these equations.

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### References

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